

Missions for Portable Meter Class Telescopes

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Light Bucket Astronomy
Conference

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Occultations by Trans-Neptunian Objects

- Certainly, observation of occultations by TNO's is a major justification for portable meter-class scopes
- TNO's move slowly so occultations are rare; if 16th and 17th-mag. stars can be considered, opportunities will be more frequent
- The larger aperture will give higher S/N for the brighter events, allowing information about TNO atmospheric fading to be measured, and/or higher time resolution of the recordings
- Large portable telescopes will allow filling in gaps in coverage by the fixed-site observatories that are now used for the fainter events
- This will result in more chords, better distributed across the object, with smaller timing errors (with shorter integration and dead times)
- More about TNO occultations have been discussed by others at this conference, so I'll talk more about other applications

Lunar Occultations

- When not making other observations, lunar occultations can be observed, if for nothing else, as practice, and to train new operators, to familiarize them with accurate timing procedures as well as with operating the large telescope
- Several events should be visible each night
- Issues with baffling and glare when the Moon is gibbous will need to be dealt with, which could help for other types of observation in twilight or light-polluted places
- With higher S/N, double stars and angular diameters might be recorded for many more fainter stars where only timings were possible with smaller scopes
- That is, Fresnel diffraction phenomena can be measured and deconvolved to reveal more information about occulted stars

Lunar Grazing Occultations

- A portable meter-class telescope can be transported into the path of a grazing occultation where multiple contacts at different position angles (PA's) can be recorded
- Video recordings with smaller telescopes should be made at locations about 20m north and south of the large telescope, so that local lunar slopes can be determined
- The deconvolution of the diffraction patterns at different PA's, as determined from the slope measurements, can reveal the light distribution across the occulted star
- This might reveal close faint companion stars, star spots, and disks of incandescent material, but probably not planets

Best 2011 N. American Lunar Grazes

RASC	2011				%	L	w.	UT	w.	end	States,	
#	Date	Star #	D	mag.	Sn1	h	m	long	lat	Prov.	Star Name	
2	JAN 1	ZC 2347		4.5	9-	N	13	32.1	-120	41 NV-TX	omicron Sco	
5	JAN 8	ZC 3320	T	5.0	18+	S	22	56.1	-92	42 IA-QC	kappa Aqr (Situla)	
6	JAN 9	ZC 3326	O	6.4	19+	S	1	3.2	-111	23 sBaja-NC		
8	JAN 10	ZC 3453	V	5.0	28+	S	2	9.7	-130	48 BC-SK	kappa Psc	
14	JAN 18	ZC 946	Q	3.5	95+	S	0	24.5	-128	40 OR-ON	eta Gem (Propus)	
15	JAN 18	ZC 976	G	2.9b	96+	S	3	54.8	-130	39 nCA-NC	mu Gem (Tejat)	
23	JAN 28	ZC 2270		5.4	31-	S	8	28.7	-72	45 VT-ME		
26	JAN 29	ZC 2442		5.9	20-	S	11	50.3	-111	39 UT-LA		
38	FEB 11	ZC 472	K	4.9	50+	N	8	22.6	-122	55 BC-SK	zeta Ari	
46	FEB 14	ZC 916	Q	4.3	79+	N	8	43.1	-130	40 sCA-MX	1 Gem	
62	MAR 13	ZC 976	G	2.9d	60+	N	23	23.7	-130	49 BC-QC	mu Gem (Tejat)	
77	MAR 25	ZC 2513	K	4.2	61-	N	11	17.8	-130	38 sCA-TX	44 Oph	
81	APR 7	ZC 486	H	5.3	10+	N	0	0.2	-80	41 PA-NJ	tau Ari	
85	APR 8	ZC 660	V	4.3	19+	N	5	22.6	-128	55 BC-WY	upsilon Tau	
86	APR 10	ZC 946	Q	3.5	36+	N	4	28.9	-130	42 CA-MX	eta Gem (Propus)	
87	APR 11	ZC 1077	B	4.0	44+	N	1	16.8	-85	55 QC-NS	zeta Gem (Mekbuda)	
132	JUL 7	ZC 1713	V	5.6	37+	N	4	37.7	-130	41 n.Calif.		
139	JUL 26	ZC 660	V	4.3	21-	N	7	48.3	-90	22 s.Fla.	upsilon Tau	
152	AUG 9	ZC 2513	K	4.2	79+	S	7	55.0	-130	36 nCA-ID	44 Oph	
153	AUG 10	ZC 2779		3.8	92+	S	23	37.7	-68	45 ME-NS	omicron Sgr	
160	AUG 22	ZC 599		4.4	46-	N	6	37.2	-106	26 MX(NL)-MA-NL	37 Tau	
168	AUG 25	ZC 1077	B	4.0d	17-	N	12	39.4	-130	31 sCA-AZ	zeta Gem (Mekbuda)	
173	SEP 4	ZC 2290	L	2.3b	42+	N	1	18.5	-130	32 MX-PA	delta Sco (Dschubba)	
194	OCT 16	ZC 656	V	4.2	85-	N	5	32.1	-113	20 TX-NC	kappa Tau	
198	OCT 17	ZC 817	T	4.9	76-	S	10	54.1	-130	46 WA-WI	114 Tau	
204	OCT 21	ZC 1341	Y	4.3	37-	S	9	17.3	-130	45 WA-NS	alpha Cnc (Acubens)	
211	NOV 5	ZC 3326	O	6.4	73+	S	4	19.7	-130	32 CA(SF)-MB		
222	NOV 28	ZC 2746		5.7	10+	S	1	23.5	-130	45 WA-AB	33 Sgr	
224	NOV 29	ZC 3015	I	5.2	25+	S	22	49.4	-91	39 MO-QC	tau Cap	
234	DEC 10	X 6484	M	8.0	0E	S	14	9.2	-130	51 BC-ID	(total lunar eclipse)	
244	DEC 18	ZC 1713	V	5.6	47-	S	7	0.9	-91	48 MN-ME		

Extra-Solar Planetary Transits

- Although extra-solar planets are unlikely to be detected from occultations (but maybe not impossible with the largest ones?),
- Transit of them across their stars can be observed, as is already being done
- The larger aperture will give better S/N for these observations, allowing detection of smaller planets than was possible with smaller telescopes
- Of course, this would be done when there aren't any occultations to chase

Asteroidal Occultations

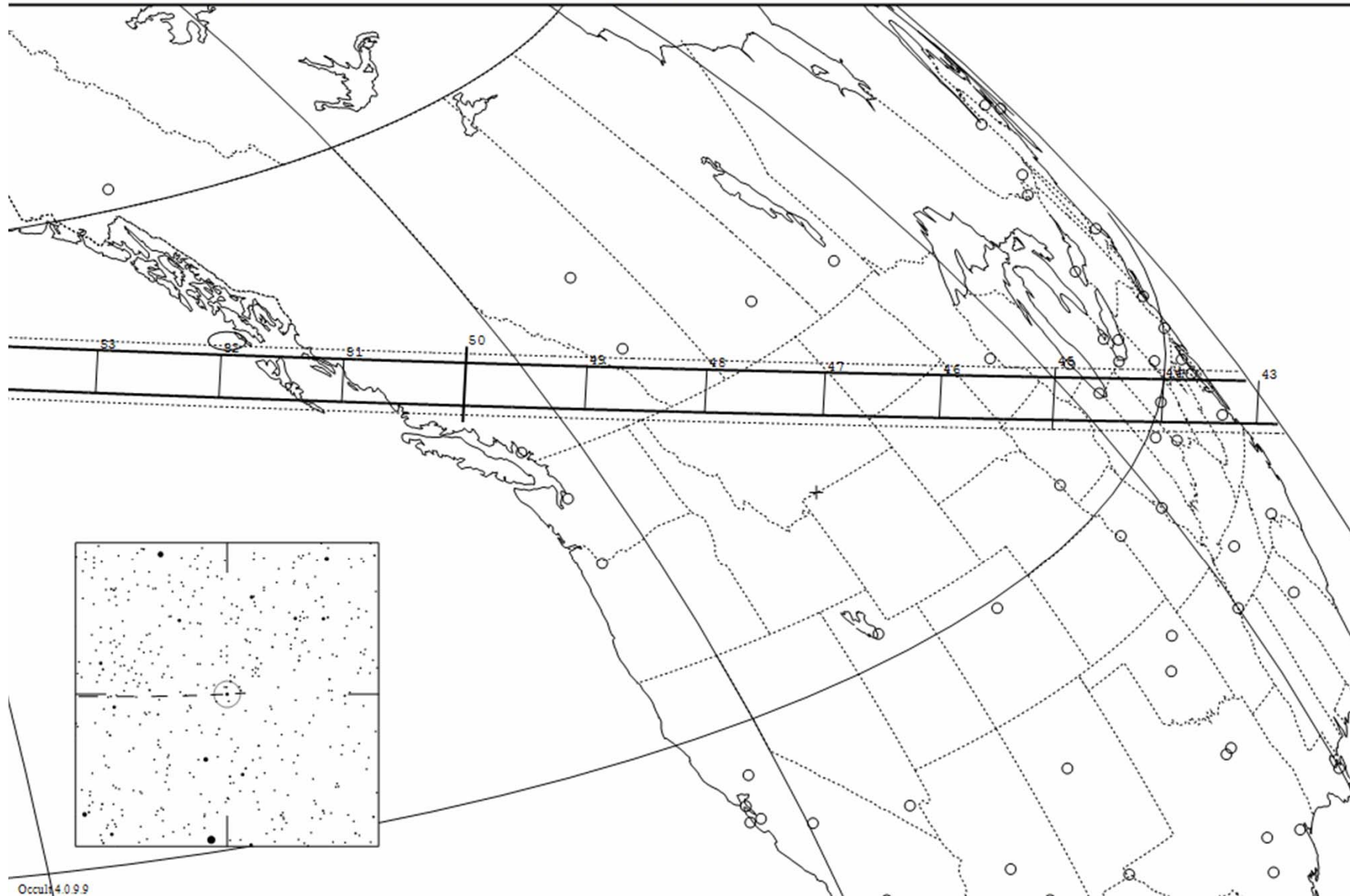
- Besides TNO's, larger portable telescopes could be used to observe occultations by more "ordinary" closer asteroids
- Besides main-belt objects, the fainter stars recordable with large scopes will allow more opportunities for other rare occultations, such as by Centaur objects and comets
- For many main-belt asteroids, whose paths can be predicted quite accurately, the larger aperture will record diffraction phenomena,
- allowing detection of close faint companion stars, star spots, and disks of incandescent material, but probably not planets, at ten or more times the resolution of lunar occultations
- The next four slides show some outstanding asteroidal occultation possibilities in coming months

419 Aurelia occults HIP 25908 on 2011 Feb 2 from 7h 43m to 8h 11m UT

Star:
Mv = 8.8 Mp = 9.0 Mr = 8.7
RA = 5 31 46.2305 (J2000)
Dec = 19 33 4.302
[of Date: 5 32 28, 19 33 32]
Prediction of 2010 Oct 20.0

Max Duration = 21.5 secs
Mag Drop = 5.0 (4.7r)
Sun : Dist = 130 deg
Moon: Dist = 139 deg
: illum = 1 %
E 0.031"x 0.015" in PA 91

Asteroid:
Mag = 13.8
Dia = 129km, 0.070"
Parallax = 3.484"
Hourly dRA = -0.831s
dDec = 0.62"

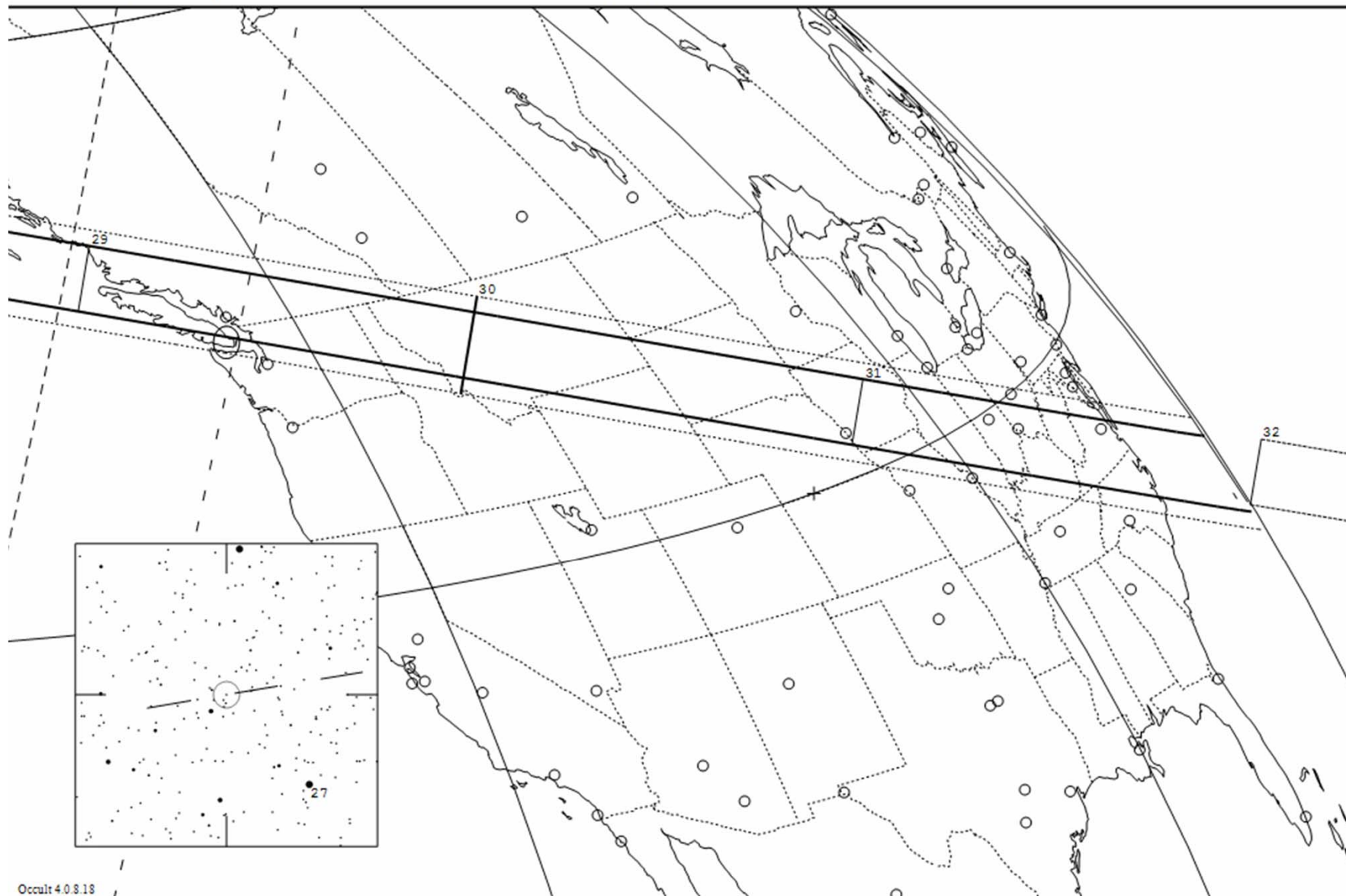


7 Iris occults TYC 0808-00566-1 on 2011 Apr 30 from 4h 24m to 4h 32m UT

Star:
Mv = 10.3 Mp = 11.1 Mr = 9.9
RA = 8 28 58.7143 (J2000)
Dec = 13 14 45.465
[of Date: 8 29 37, 13 12 22]
Prediction of 2010 May 17.0

Max Duration = 10.1 secs
Mag Drop = 0.8 (0.7r)
Sun : Dist = 87 deg
Moon: Dist = 121 deg
: illum = 9 %
E 0.030"x 0.024" in PA 1

Asteroid:
Mag = 10.3
Dia = 200km, 0.123"
Parallax = 3.912"
Hourly dRA = 2.956s
dDec = -7.22"

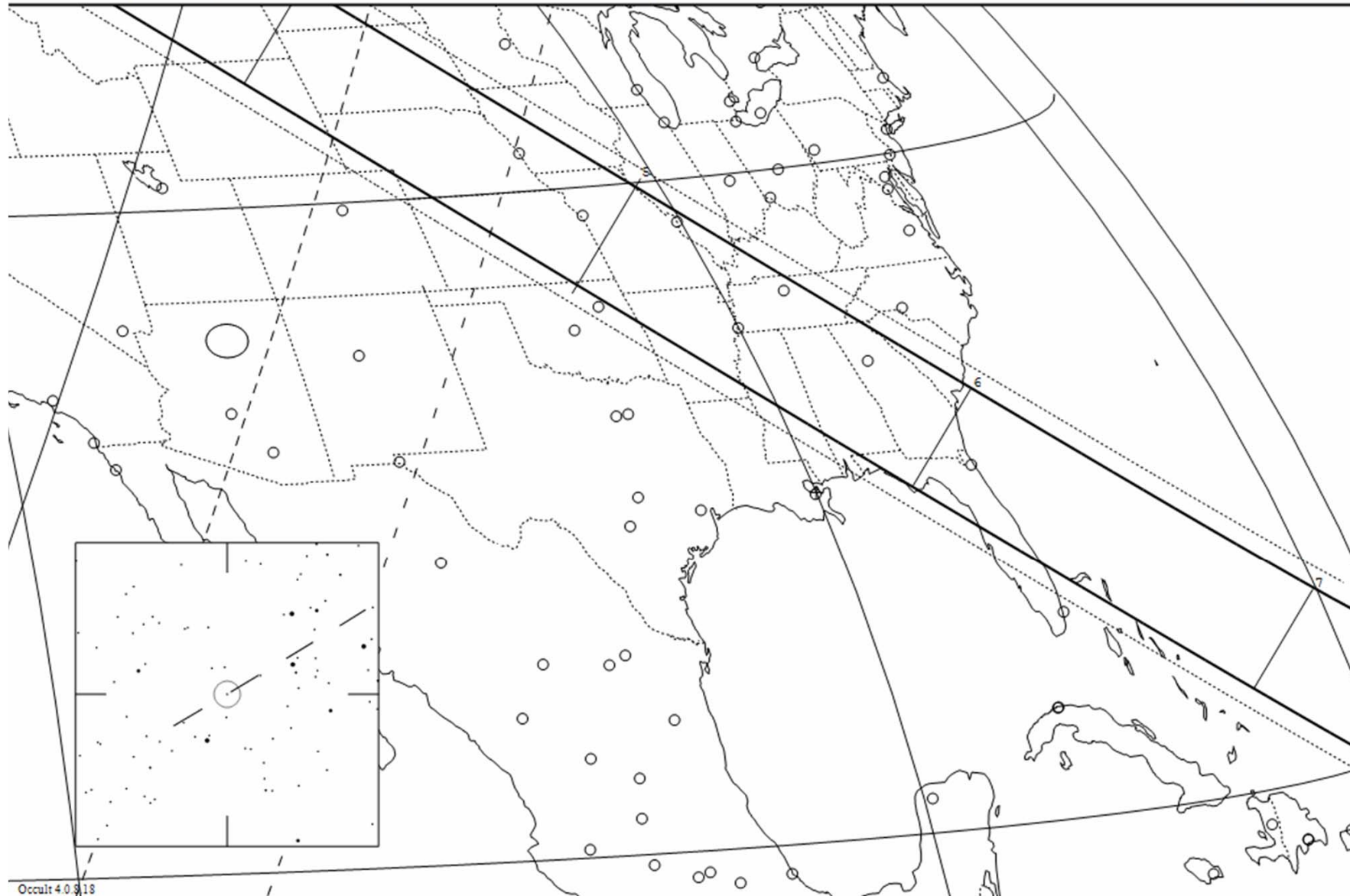


52 Europa occults TYC 0292-00339-1 on 2011 Jul 4 from 3h 0m to 3h 8m UT

Star:
Mv = 10.1 Mp = 10.3 Mr = 10.0
RA = 12 33 5.2376 (J2000)
Dec = 4 31 28.510
[of Date: 12 33 41, 4 27 37]
Prediction of 2010 May 17.0

Max Duration = 17.3 secs
Mag Drop = 2.2 (1.9r)
Sun : Dist = 84 deg
Moon: Dist = 50 deg
: illum = 9 %
E 0.029"x 0.023" in PA 93

Asteroid:
Mag = 12.1
Dia = 341km, 0.156"
Parallax = 2.924"
Hourly dRA = 1.869s
dDec = -16.62"



Occultation of ZC 3339 by the binary asteroid (90) Antiope, 2011 July 19 morning

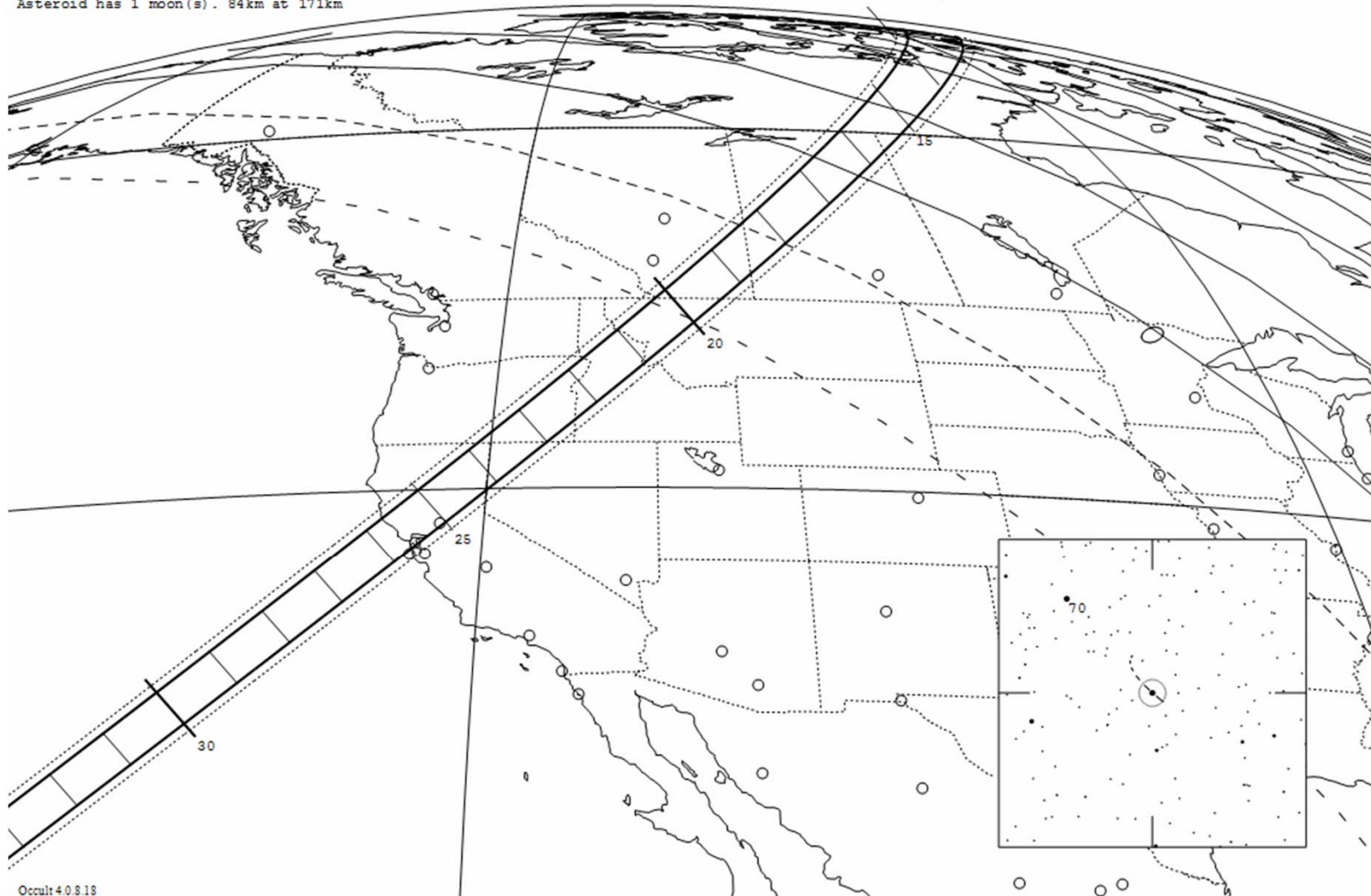
90 Antiope occults HIP 112420 on 2011 Jul 19 from 10h 13m to 11h 6m UT

Star:
Mv = 6.7 Mp = 8.3 Mr = 5.8
RA = 22 46 14.2134 (J2000)
Dec = -11 9 59.057
[of Date: 22 46 53, -11 6 6]
Prediction of 2010 May 17.0

Max Duration = 39.5 secs
Mag Drop = 5.8 (6.2r)
Sun : Dist = 138 deg
Moon: Dist = 10 deg
: illum = 83 %
E 0.025"x 0.016" in PA 72

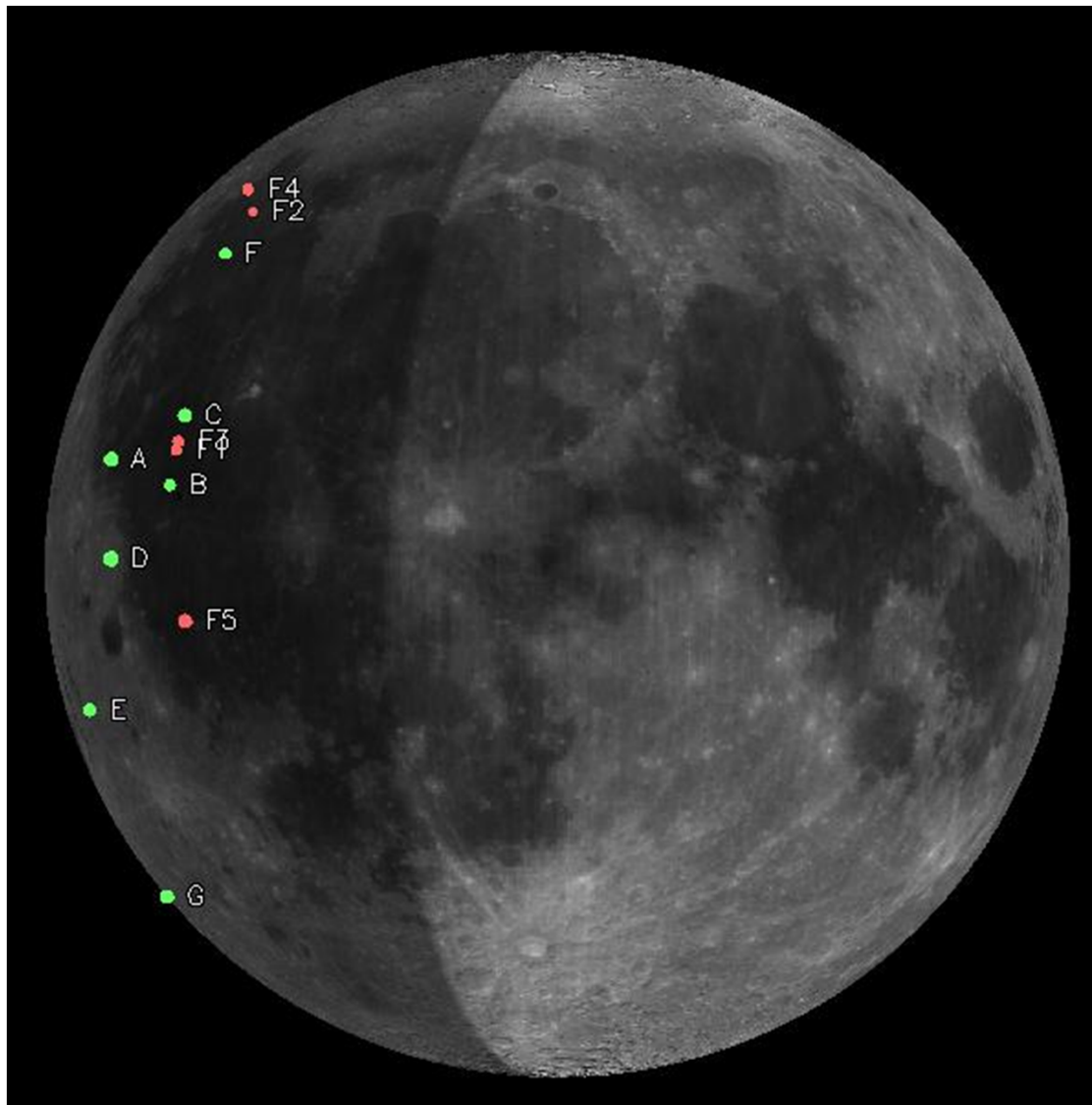
Asteroid:
Mag = 12.5
Dia = 120km, 0.091"
Parallax = 4.811"
Hourly dRA = -0.421s
dDec = -5.47"

Asteroid has 1 moon(s). 84km at 171km



Locations of 1999 Leonids Lunar Impacts

that I recorded Nov. 18 with 13cm telescope at Mt. Airy, MD; many fainter ones can be recorded with 1m scopes



The green dots mark impact flashes that were confirmed by other IOTA observers at distant locations.

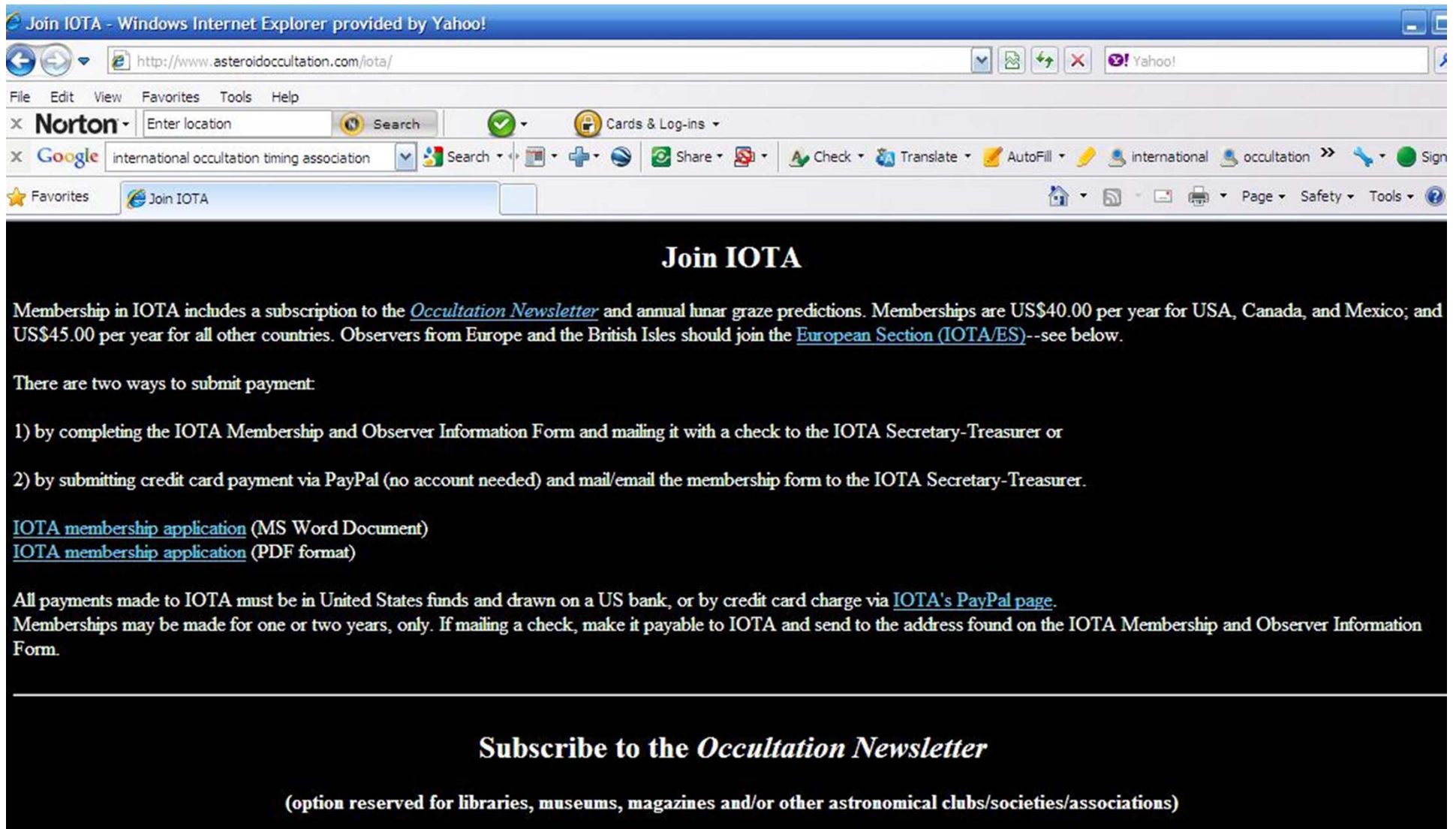
This is another activity where larger telescopes would be useful; with diffraction gratings and/or higher-speed recordings, more information about these very transient events could be learned.

Leonid Lunar Impact

Recorded 2001 Nov. 18, 23:19:15 UT, Laurel, MD
confirmed by Tony Cook at Arlington, VA
and Roger Venable, Augusta, GA



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